

## DIAGNOSTICS OF NITROGEN-METHANE ATMOSPHERIC GLOW DISCHARGE USED FOR A MIMIC OF PREBIOTIC ATMOSPHERE

V. MAZÁNKOVÁ<sup>a,\*</sup>, L. TÖRÖKOVÁ<sup>a</sup>, D. TRUNEC<sup>b</sup>, F. KRČMA<sup>a</sup>, S. MATEJČÍK<sup>c</sup>,  
N. J. MASON<sup>d</sup>

<sup>a</sup> Faculty of Chemistry, Brno University of Technology, Purkyňova 118, 612 00 Brno, Czech Republic

<sup>b</sup> Faculty of Science, Masaryk University, Kotlářská 2, 611 37 Brno, Czech Republic

<sup>c</sup> Faculty of Mathematics, Physics and Informatics, Comenius University, 842 48 Bratislava, Slovakia

<sup>d</sup> Department of Physical Sciences, Open University, Walton Hall, Milton Keynes MK7 6AA, UK

\* mazankova@fch.vut.cz

**Abstract.** The exploration of planetary atmosphere is being advanced by the exciting results of the Cassin-Huygens mission to Titan. The complex chemistry revealed in such atmospheres leading to the synthesis of bigger molecules is providing new insights into our understanding of how life on Earth developed. This work extends our previous investigation of nitrogen-methane (N<sub>2</sub>-CH<sub>4</sub>) atmospheric glow discharge for simulation chemical processes in prebiotic atmospheres. In presented experiments 2% of water vapor were added to nitrogen-methane gas mixture. Exhaust products of discharge in this gas mixture were in-situ analysed by Fourier Transform Infra Red spectroscopy (FTIR). The major products identified in spectra were: hydrogen cyanide, acetylene and acetonitrile.

**Keywords:** prebiotic atmosphere, Fourier Transform Infra Red spectroscopy, glow discharge.

### 1. Introduction

More than 60 years ago, S. Miller and H. Urey [1], in their exciting work, demonstrated the synthesis of several amino acids from a mixture of reducing gases (CH<sub>4</sub>, NH<sub>3</sub>, H<sub>2</sub>O, and H<sub>2</sub>) treated with electric discharge. The following explorations showed that a broad array of amino acids could be synthesized, but there was no evidence that all of the fundamental molecules of the RNA genetic code could be produced alongside others in this type of experiment [2, 3]. Recently it was also referred to a positive result on qualitative detection of RNA nucleobases and manifold amino acids from tholines created in a N<sub>2</sub>, CH<sub>4</sub>, CO mixture [4]. This experiment simulated the atmosphere of Titan by electric discharge. Such experimental results as well as theoretical expectations [5] show that reduced, relatively reactive atmospheres are likely to be more efficient for the synthesis of biomolecules [6]. The gliding arc configuration of atmospheric pressure discharge has been shown to be a good mimic of processes in the prebiotic atmospheres [7] being used to replicate physical and chemical conditions on Titan. It is the only lunar body with significant quantities of methane (CH<sub>4</sub>) and nitrogen (N<sub>2</sub>) in its atmosphere [8, 9]. The chemical composition of the Titan's atmosphere is considered to be similar to the atmosphere of early Earth and it is favorable for formation of complex molecules containing C, N and H. The solar ultraviolet radiation and magnetospheric electrons are responsible for the generation of primary radicals and other neutral species, which initiate chains of chemical reactions that finally result in the formation of various organic molecules in the Titan atmosphere.

This makes Titan as planetary-scale laboratory for the synthesis of complex organic molecules [10]. The composition of early Earth atmosphere was deeply discussed in many studies [11–13].

The present work is focused on the experimental study of gaseous products produced in the atmospheric pressure glow discharge fed by N<sub>2</sub>, CH<sub>4</sub> gas mixtures with CH<sub>4</sub> concentrations in the range from 1% to 5% and admixture of 2% water vapor (H<sub>2</sub>O). The atmospheric DC glow discharge is a source of electrons, ions and neutral radicals. All these species initiate a complex chemical processes under laboratory conditions, which are similar to the processes which may occur in the prebiotic atmosphere. The neutral products generated in the discharge were identified and quantified by the means of the Fourier-Transform-Infra-Red spectroscopy (FTIR). In more detail the influence of the H<sub>2</sub>O admixture on production of HCN (the major neutral product) and on the formation of NH<sub>3</sub> has been studied.

### 2. Experimental setup

The experimental setup schematic drawing is shown in Figure 1. The special high vacuum stainless steel reactor was constructed for our experiments to prevent any oxygen contamination during the experiments. Nitrogen and methane flows were automatically controlled by Bronkhorst controllers. The measurements were carried out at total gas flows 100 sccm at atmospheric pressure and laboratory temperature. The discharge electrode system had the standard configuration of the gliding arc discharge. The discharge was formed in the stable abnormal glow regime, and

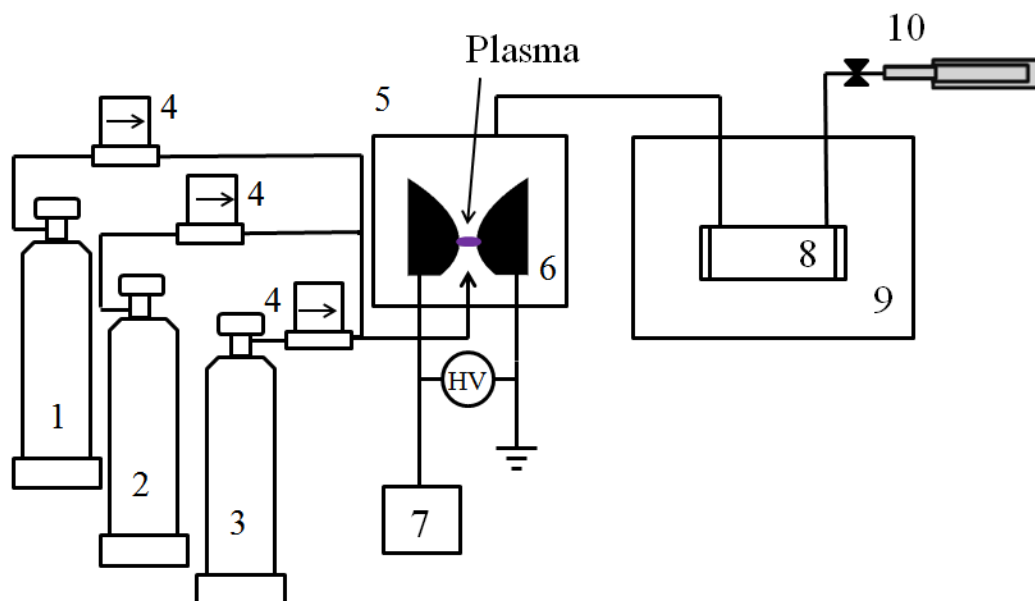


Figure 1. Experimental setup. 1- storage bottle of nitrogen, 2- storage bottle of methane, 3- storage bottle of carbon dioxide, 4- Bronkhorst controllers, 5- reactor body, 6- electrode system, 7- oscilloscope, 8- IR gas cell, 9- FTIR spectrometer, 10- exhaust.

plasma occurred between the electrodes at their shortest distance of 2 mm in the form of a plasma channel of 1 mm in its diameter. The reactor chamber volume was 0.5 l. The discharge was supplied by a DC stabilized HV source. Discharge breakdown voltage was 1500 V, a stable plasma channel was operating at 400 V at current in range 15 mA to 40 mA during all presented experiments. The measurements were performed for different  $N_2:CH_4$  ratios in the range from 1 % to 5 % of methane in nitrogen (both gases having quoted purities of 99.995 %). There was connected bottle gas washing with a high purity water just before the entrance to the reactor. The amount of added water vapour was estimated at 2 % in gas mixture in all performed experiments. The exhaust gas was analyzed in-situ by FTIR spectroscopy using IR cell with total length of 15 cm.

The electrode system is made of stainless steel and it is necessary to clean it after each set of experiments. The deposit created during discharge operation is shown in Figure 2. Before starting the experiments, the discharge chamber was pumped down to 1 Pa for 1 h and then was filled with the investigated gas mixture up to the pressure of 101 kPa. Atmospheric pressure during the experiments was maintained by a slight pumping through the needle valve.

### 3. Results and Discussion

A typical FTIR spectrum showing the products formed in the nitrogen discharge fed by 4 % of  $CH_4$  and 2 % of  $H_2O$  is shown in Figure 3. Similar spectra were also observed for other  $N_2:CH_4:H_2O$  ratios. Hydrogen cyanide (HCN) was found to be the most abundant product at wavenumbers of  $1430\text{ cm}^{-1}$  and of



Figure 2. Photos of the electrodes with deposited layers.

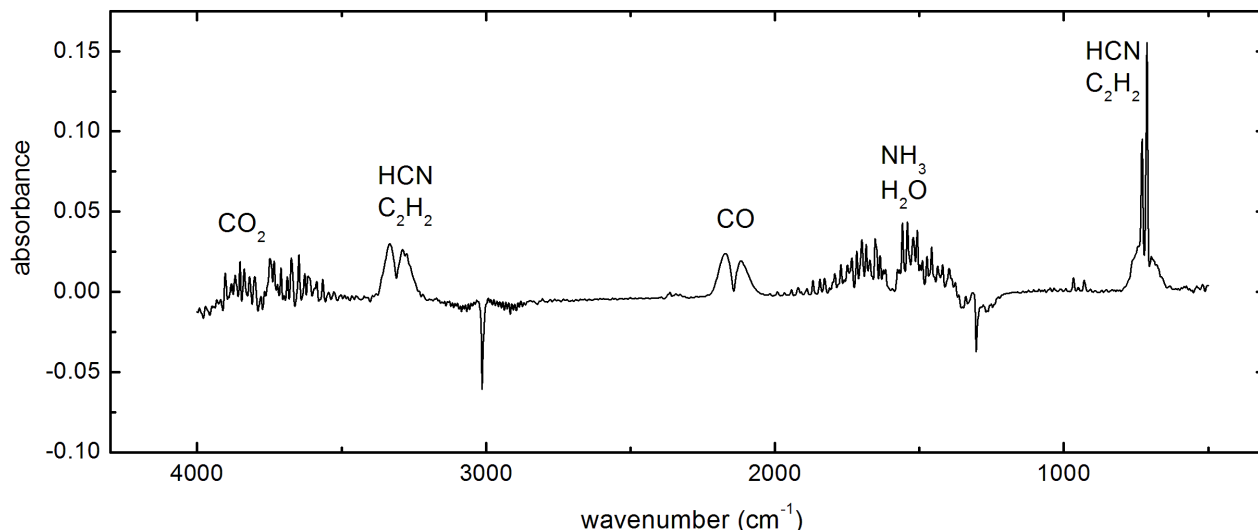


Figure 3. FTIR spectrum recorded in mixture  $N_2:CH_4 = 96:4$  with 2%  $H_2O$  vapor addition at flow rate of 100 sccm and discharge current of 20 mA.

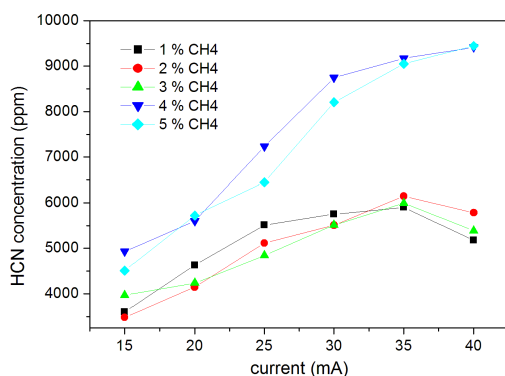


Figure 4. The dependence of hydrogen cyanide concentration on discharge current for different initial concentration of methane.

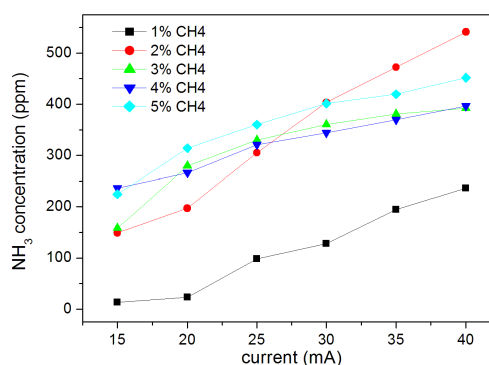


Figure 5. The dependence of ammonia concentration on discharge current for different initial concentration of methane and carbon dioxide.

$720\text{ cm}^{-1}$ . This is in agreement with [7]. Ammonia ( $NH_3$ ), which was identified at  $966\text{ cm}^{-1}$ , was also detected in our previous experiments. The other major products were acetylene ( $C_2H_2$ ) as well as carbon monoxide (CO). These products were recognized in all investigated gas mixtures. The products concentrations depend on the composition of the gas mixtures. Quantitative analysis of selected hydrocarbons were done in dependence on the methane concentration. The absorbance was calculated as an area under the recorded peaks and the concentrations of selected products were then calculated. Figures 4 and 5 show dependences of HCN ( $1430\text{ cm}^{-1}$ ) and  $NH_3$  ( $966\text{ cm}^{-1}$ ) concentrations formed under different experimental conditions. The increase of the initial discharge current from 15 mA to 40 mA leads to increase in the product yield of these compounds.

## 4. Conclusions

The gaseous products formed in the atmospheric glow discharge fed by different mixtures of nitrogen, methane and water were determined by in situ FTIR analysis. The discharge was operated in the flowing regime at different discharge currents at laboratory temperature. An in-situ FTIR technique for the exhaust gas phase sampling was successfully used for chemical analysis to deduce the gas composition in the  $N_2:CH_4:CO_2$  reactive gas mixture mimics of Titan's atmosphere. Various hydrocarbons were observed in all experiments. HCN was identified as the major gas phase product in all of measurements. Others minor products detected were  $C_2H_2$ ,  $NH_3$ ,  $CO_2$ , CO. These results are consistent with the Titan's atmospheric composition because the same compounds were detected during the Cassini-Huygens space mission. CO could have an effect on the atmospheric reactivity

of Titan. The formation of organic molecules incorporating oxygen in gases could occur in the upper atmosphere of Titan where the dissociations of N<sub>2</sub> and CO<sub>2</sub> by VUV photons and magnetospheric electrons are possible. This fact clearly demonstrates that laboratory experiments can be used for prediction of both the presence and possible concentrations of compounds which have not been detected, yet. These simple organics should be tracers of the chemical groups constituting the dusty products.

### Acknowledgements

The research was supported by Czech Ministry of Education, Youth and Sports, projects No. COST CZ LD15010 and LD15011 within the collaboration of the COST Actions CM 1401 and TD 1308 and CEEPUS network AT-0063.

### References

- [1] S. L. Miler. A production of amino acids under possible primitive earth conditions. *Science*, 117(3046):528–529, 1953. doi:10.1126/science.117.3046.528.
- [2] T. M. McCollom. Miller-Urey and beyond: What have we learned about prebiotic organic synthesis reactions in the past 60 years?. *Annu. Rev. Earth. Planet. Sci.*, 41:207–229, 2013. doi:10.1146/annurev-earth-040610-133457.
- [3] A. Johnson, H. J. Cleaves, J. L. Bada, and A. Lazcano. The diversity of the original pre-biotic soup: Re-analyzing the original Miller-Urey spark discharge experiments. *Orig. Life. Evol. Biosph.*, 39:240–241, 2009.
- [4] S. M. Hörst. Formation of amino acids and nucleotide bases in a titan atmosphere simulation experiment. *Astrobiology*, 12:809–817, 2009. doi:10.1089/ast.2011.0623.
- [5] C. Chyba and C. Sagan. Endogenous production, exogenous delivery and impact shock synthesis of organic molecules: An inventory for the origins of life. *Nature*, 355:125–132, 1992. doi:10.1038/355125a0.
- [6] C. F. Chyba, P. J. Thomas, Brookshaw, and C. L., Sagan. Cometary delivery of organic molecules to the early earth. *Science*, 249:366–373, 1990. doi:10.1126/science.11538074.
- [7] L. Torokova, J. Watson, F. Krcma, V. Mazankova, N.J. Mason, G. Horvath, and S. Matejcik. Gas chromatography analysis of discharge products in N<sub>2</sub>-CH<sub>4</sub> gas mixture at atmospheric pressure: study of mimic Titan’s atmosphere. *Contrib Plasma Phys*, 55:470–480, 2015. doi:10.1002/ctpp.201400052.
- [8] A. Coustenis, D.E. Jennings, C.A. Nixon, R.K. Achterberg, P. Lavvas, S. Vinatier, N.A. Teanby, G.L. Bjoraker, R.C. Carlson, L. Piani, G. Bampasidis, F.M. Flasar, and P.N. Romani. Titan trace gaseous composition from CIRS at the end of the Cassini-Huygens prime mission. *Icarus*, 207:461–476, 2010. doi:10.1016/j.icarus.2009.11.027.
- [9] S.M. Horst and M.A. Tolbert. The effect of carbon monoxide on planetary haze formation. *Astrophys J*, 781(53), 2014. doi:10.1088/0004-637X/781/1/53.
- [10] F. Raulin, C. Brasse, O. Poch, and P. Coll. Prebiotic-like chemistry on Titan. *Chem Soc Rev*, 41:5380–5393, 2012. doi:10.1039/C2CS35014A.
- [11] J.F. Kasting and M.T. Howard. Atmospheric composition and climate on the early earth. *Philos Trans R Soc Lond B Biol Sci*, 361:1733–1741, 2006. doi:10.1098/rstb.2006.1902.
- [12] J.F. Kasting and S. Ono. Palaeoclimates: the first two billion years. *Philos Trans R Soc Lond B Biol Sci*, 361:917–929, 2006. doi:10.1098/rstb.2006.1839.
- [13] S.L. Olson, L.R. Kump, and J.F. Kasting. Quantifying the areal extent and dissolved oxygen concentrations of archean oxygen oases. *Chem Geol*, 362:34–43, 2013. doi:10.1016/j.chemgeo.2013.08.012.